

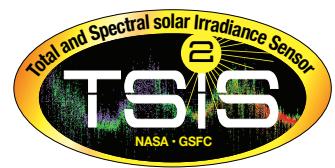


# Non-Gaussian PDFs of TOA SW Flux from MISR and CERES

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2. NASA Goddard Space Flight Center, Greenbelt, Maryland

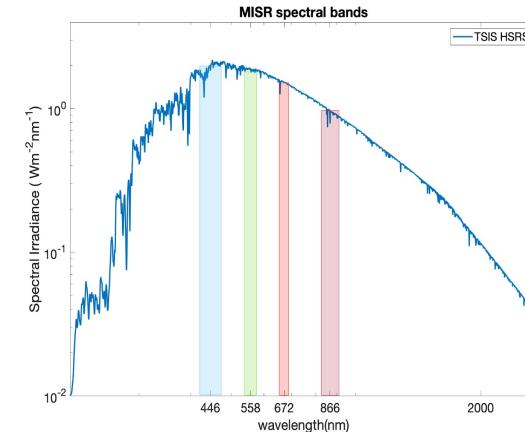
Acknowledgments : NASA Sun Climate Research



# MISR and CERES TOA SW Flux Data

## MISR (Multi-angle Imaging SpectroRadiometer) (ATBD)

- Near-simultaneous multi-angular measurements
  - Narrow-to-broad band conversion (Sun et al., 2006)
  - broad band albedo => SW Flux by  $S_0$
- 
- 275-m pixel
  - Narrow (~400km) swath
  - Scene type classification for albedo models (i.e., land, snow, ice, clouds): 1000's



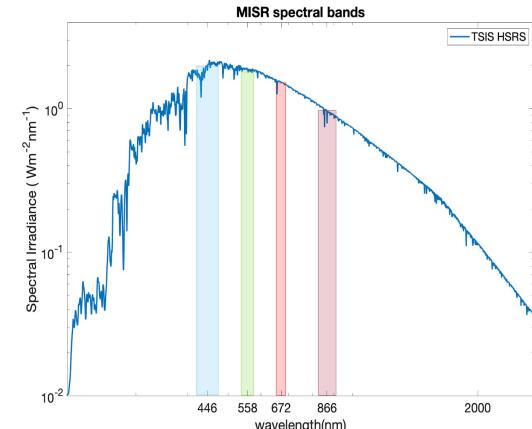
## CERES (ATBD) : Single Scanner Footprint (SSF) product

- Climatological angular distribution models (ADMs)
- Direct broad-band SW flux divided
- 10-km pixel
- Wide (~3000km) swath
- Scene type classification for ADMs: 1000's

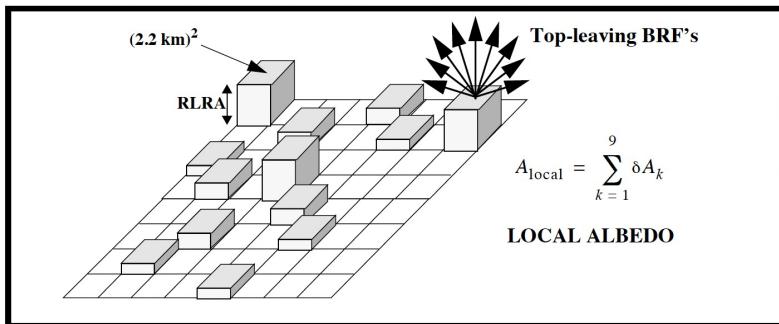
# MISR and CERES TOA SW Flux Data

## MISR (Multi-angle Imaging SpectroRadiometer) (ATBD)

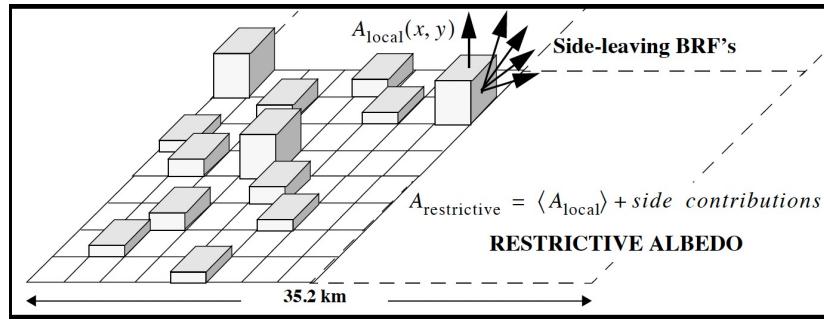
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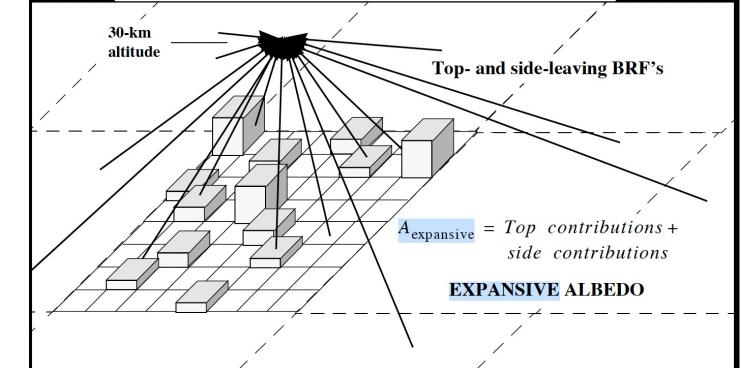
### Local Albedo



### Restrictive Albedo

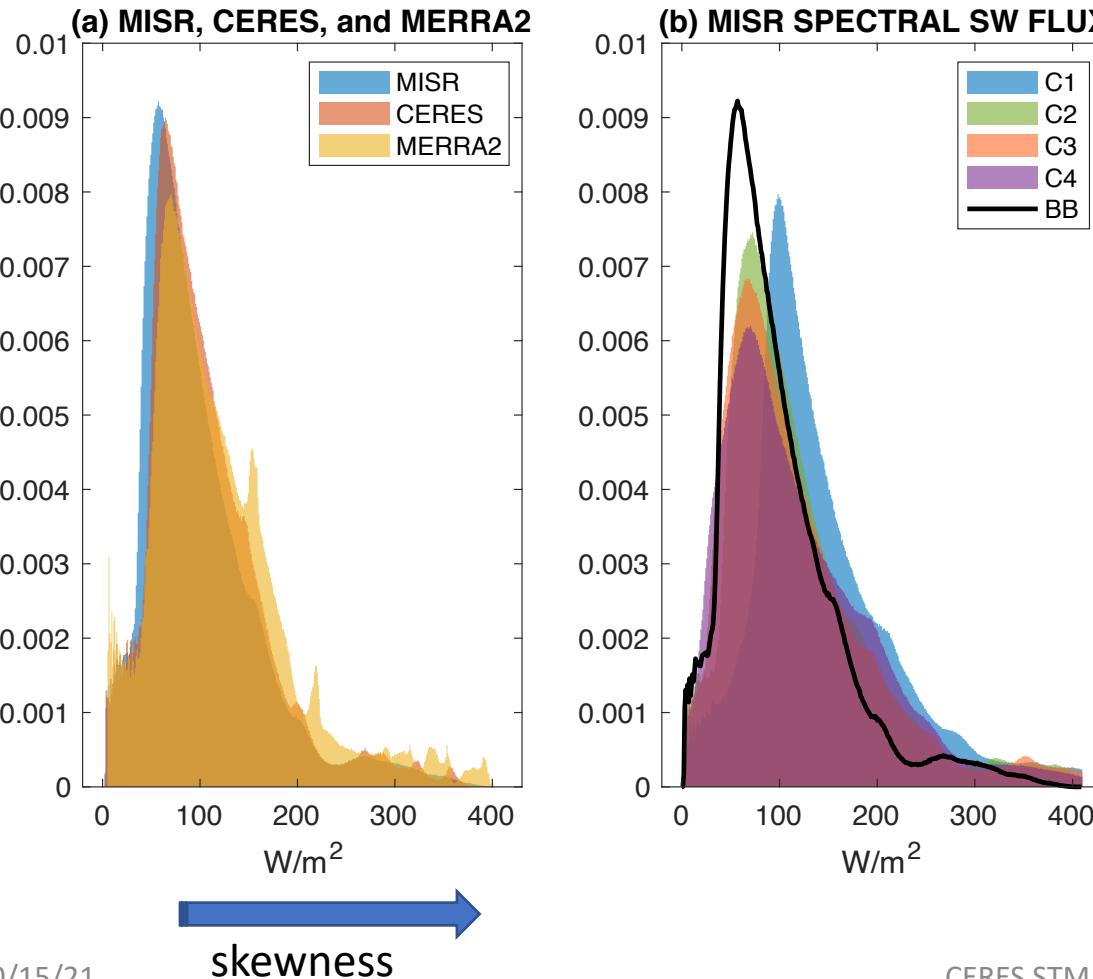


### Expansive Albedo



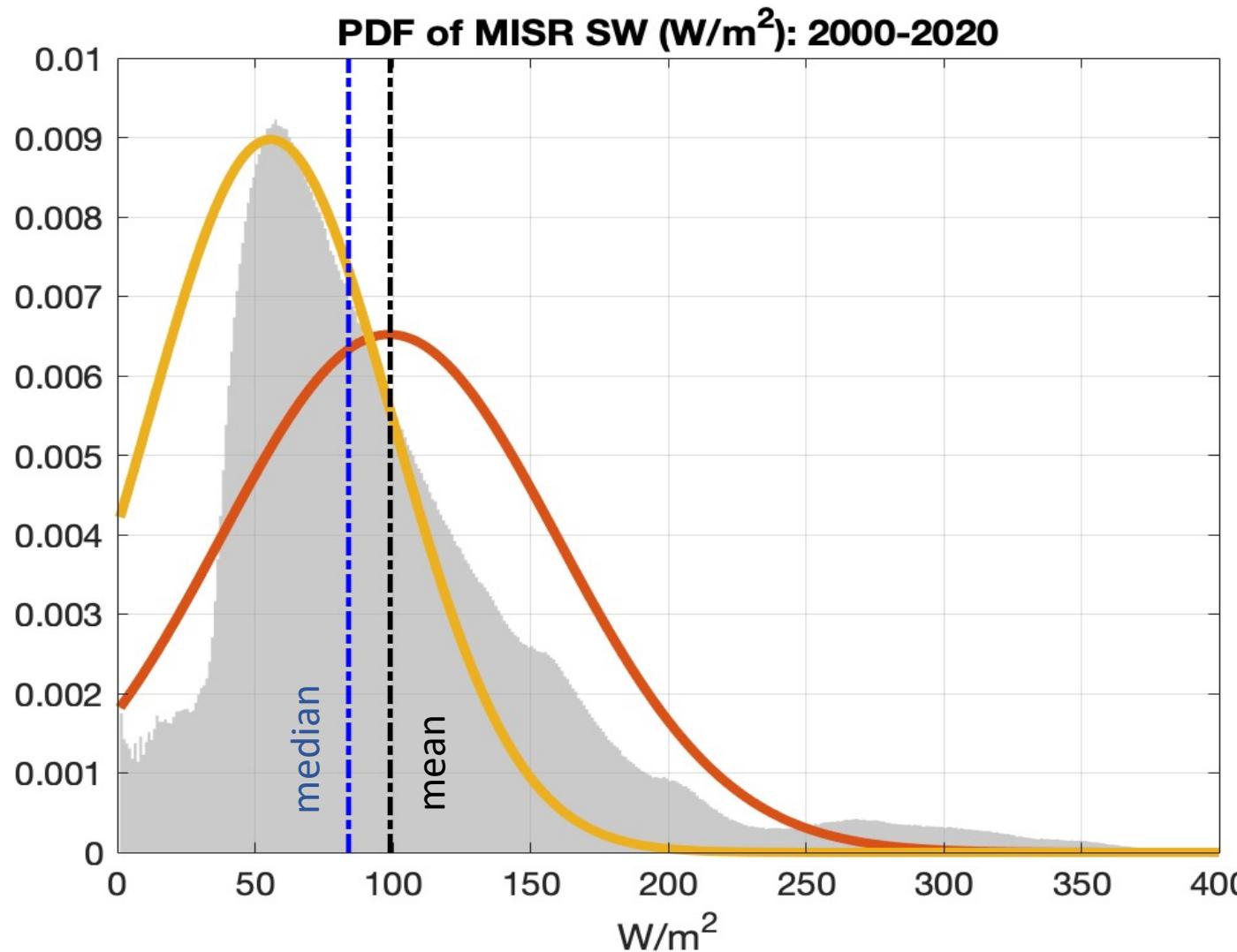
# Skewness in the TOA SW distribution

PDFs of GLOBAL ALL-SKY TOA SW : 2000-2020

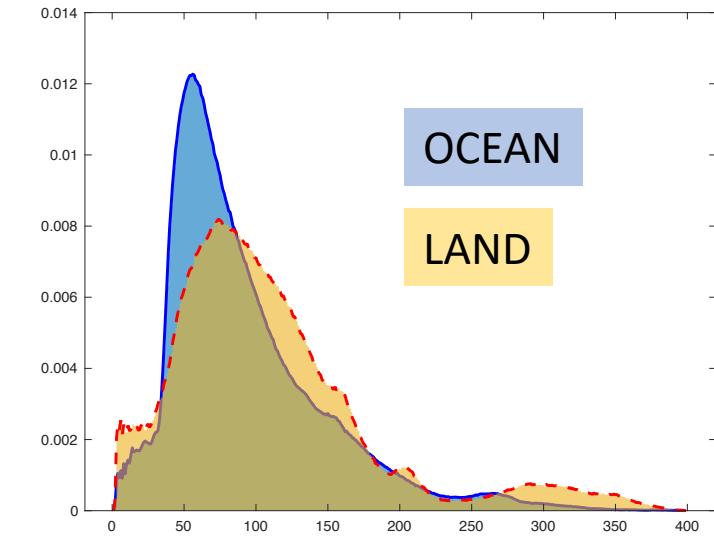


- Deviation from normal distribution
- Asymmetric from the maximum population
- Positive skewness → more data on right side of the maximum population
- One major peak near  $60\text{W/m}^2$ , MISR peak value is  $\sim 7\text{W/m}^2$  less than that of CERES
- All four of MISR spectral SW flux are showing similar positive skewness

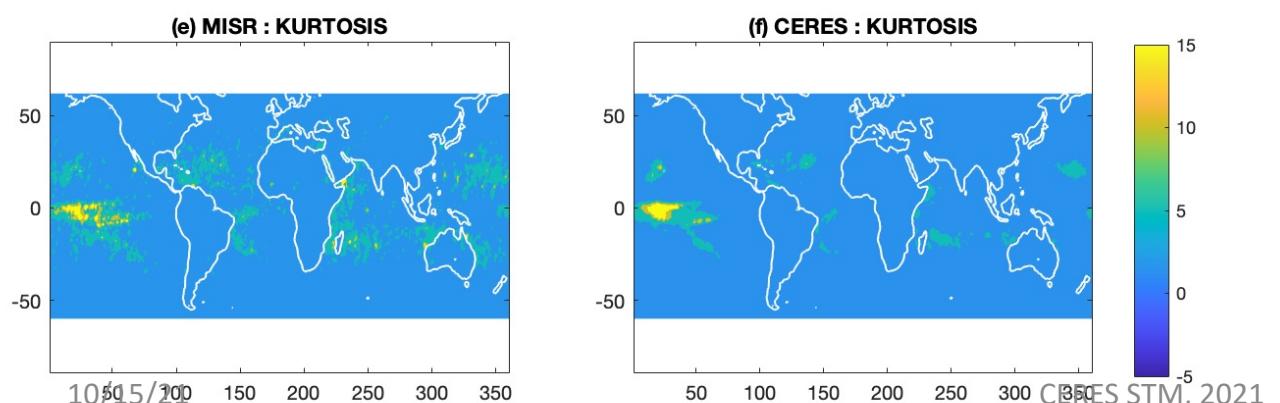
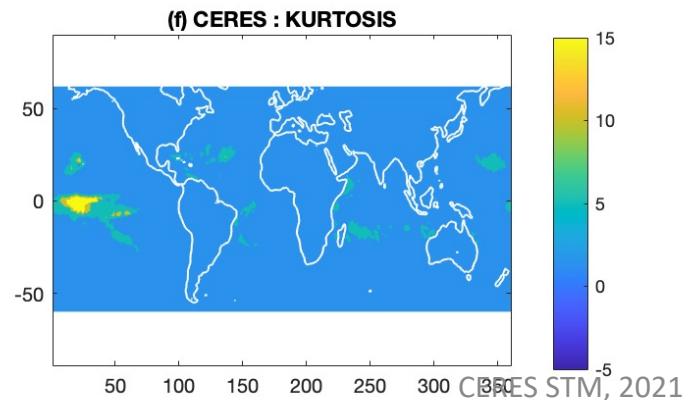
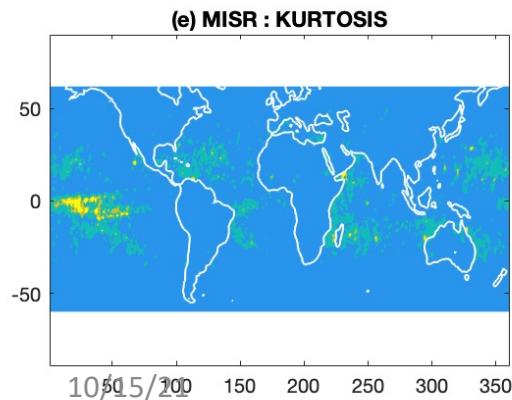
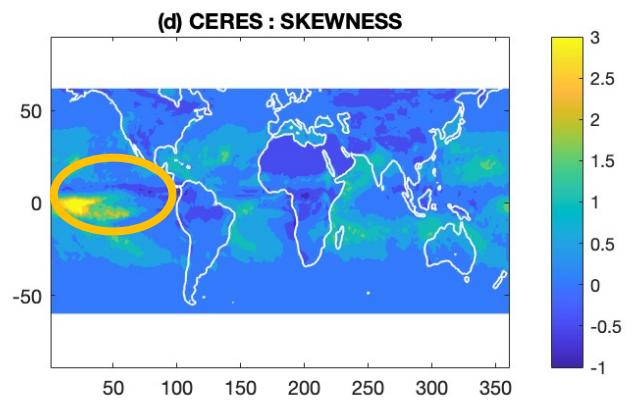
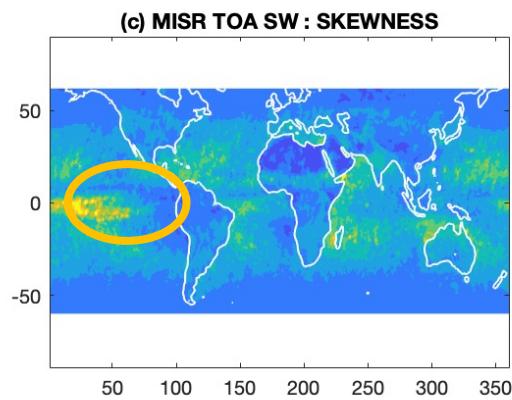
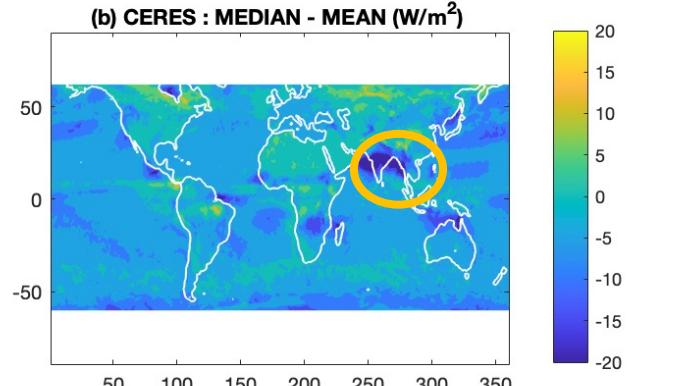
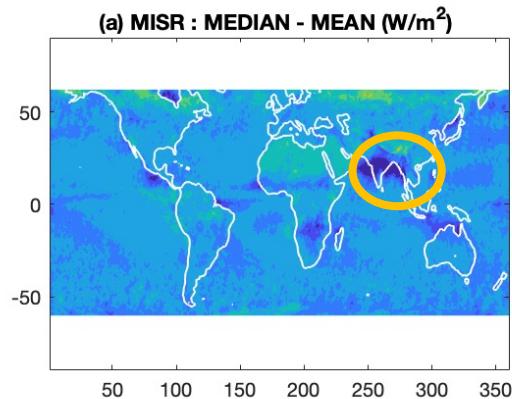
# Implications of Non- Gaussian distribution?



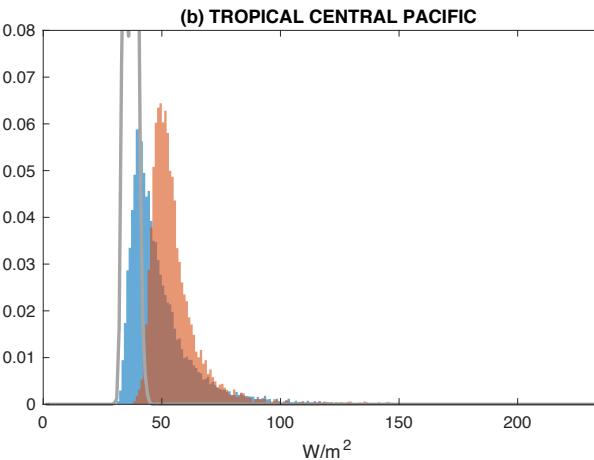
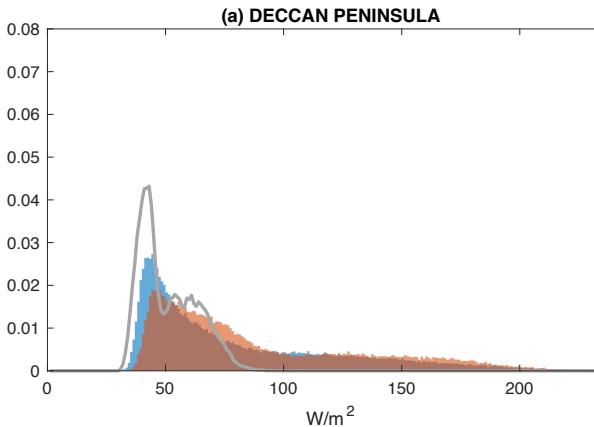
1. Deviations from mean
2. Outliers in high flux values above 300  $\text{W/m}^2$ ?



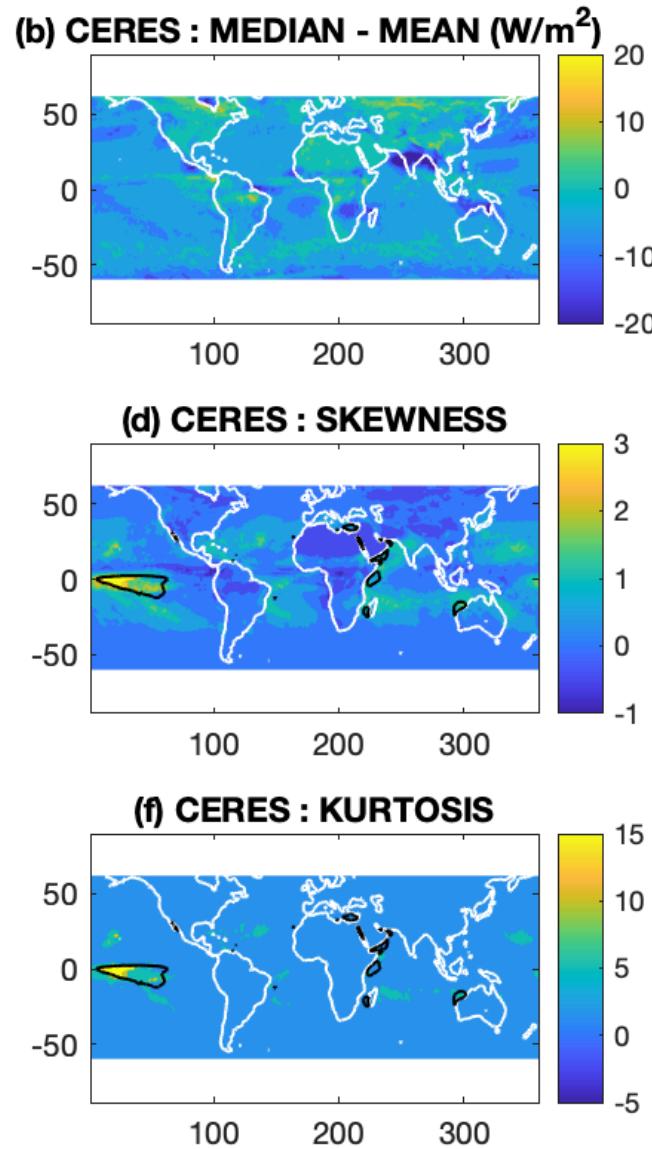
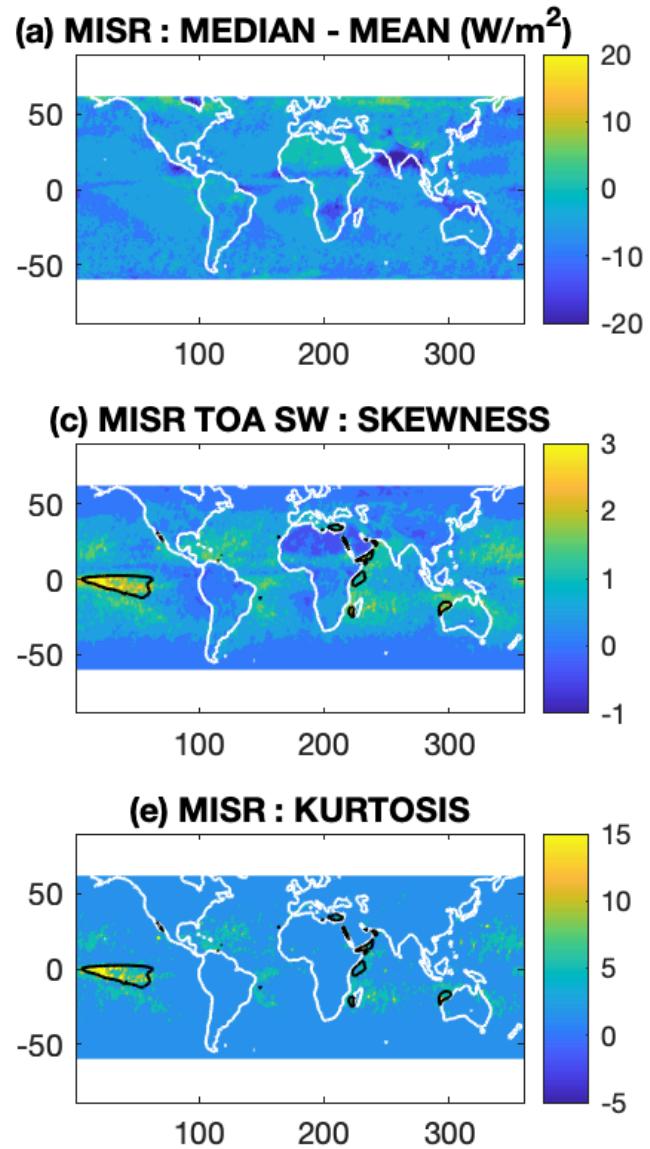
Where are the most skewed regions?



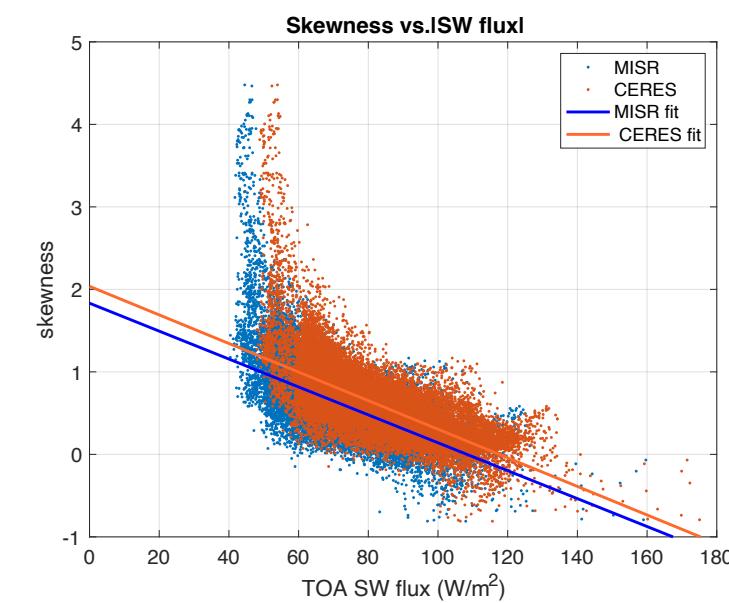
PDF of TOA SW FLUX ( $\text{W/m}^2$ )



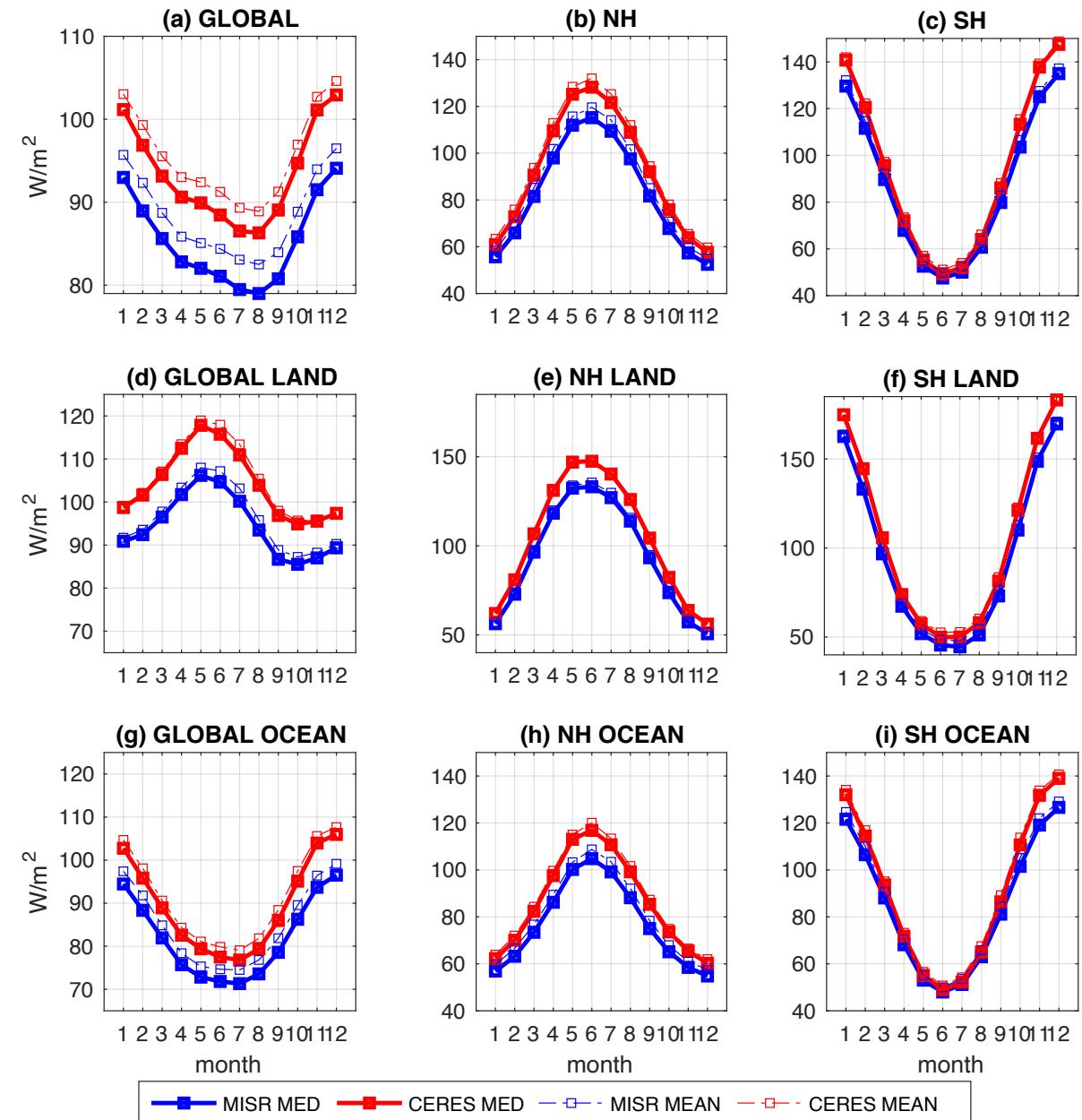
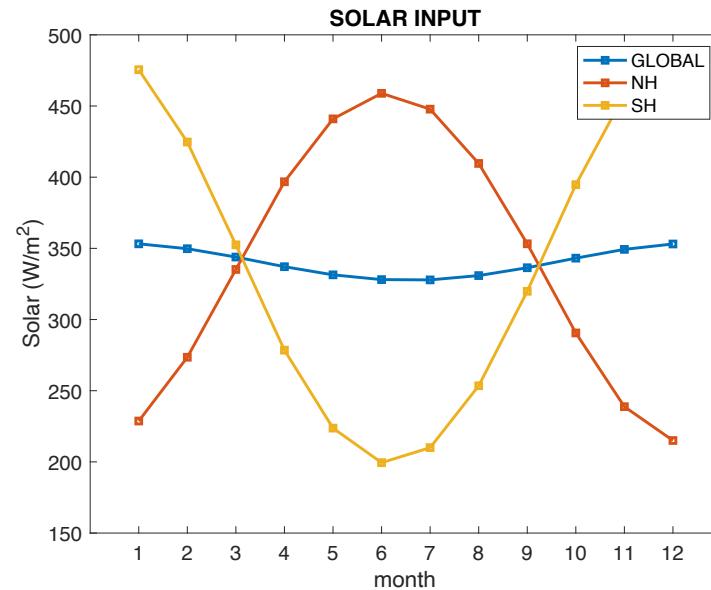
MISR    CERES    CERES CLR SKY



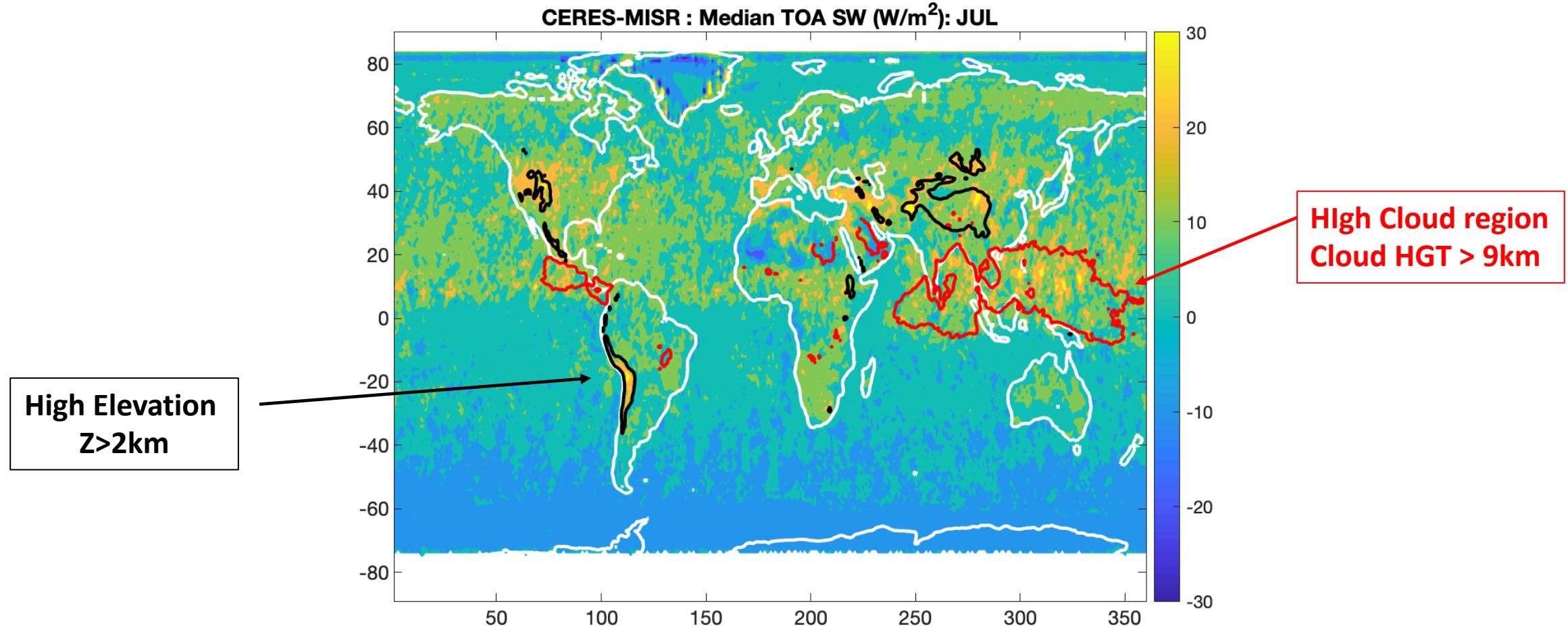
Where are the most skewed regions?



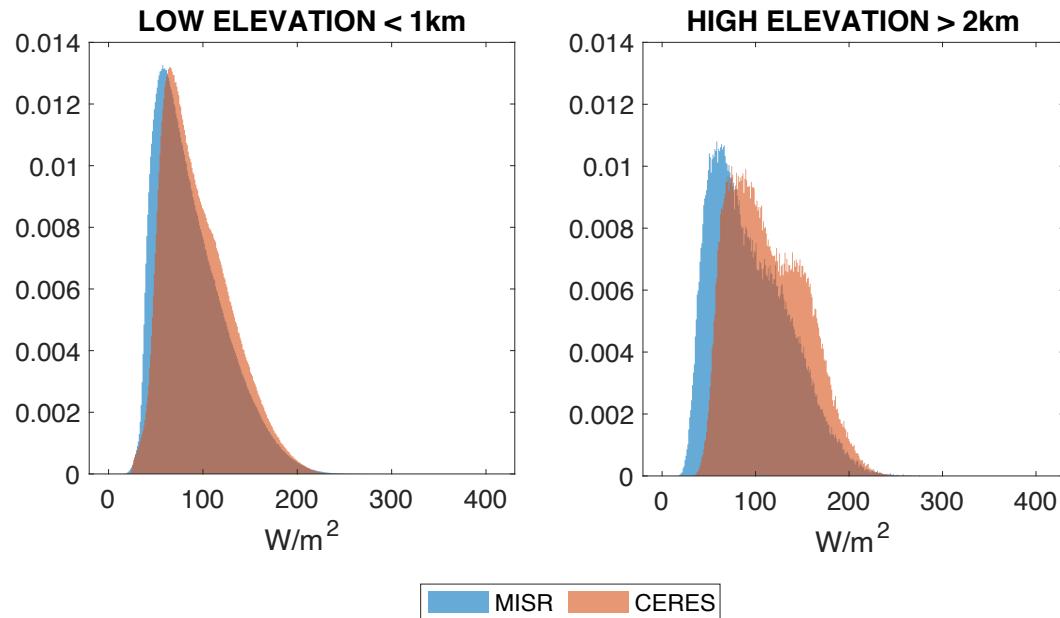
# Seasonal variation of TOA SW (2000-2020)



# Difference between CERES - MISR



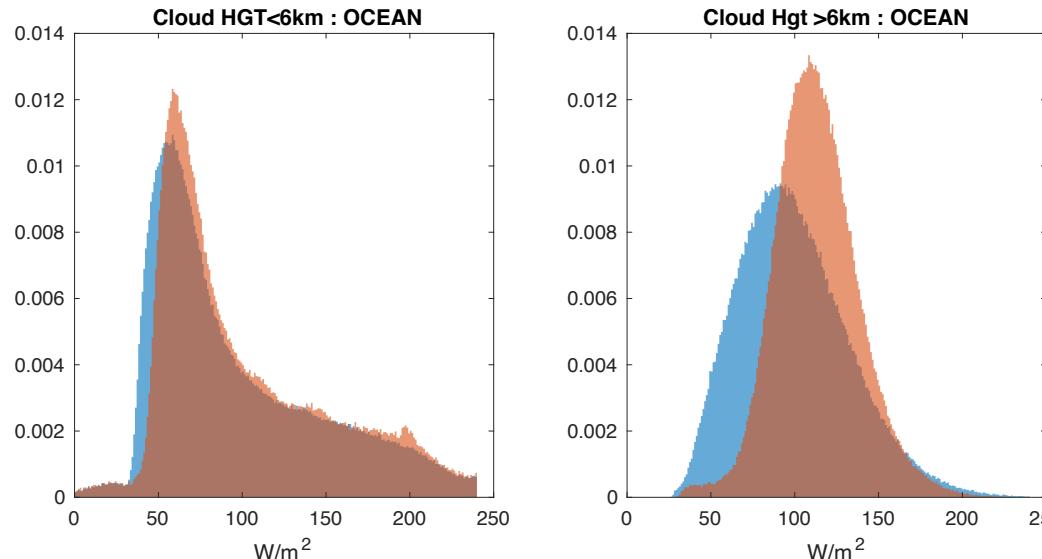
TOA SW ( $\text{W/m}^2$ ) : 60S-60N



# High vs. Low

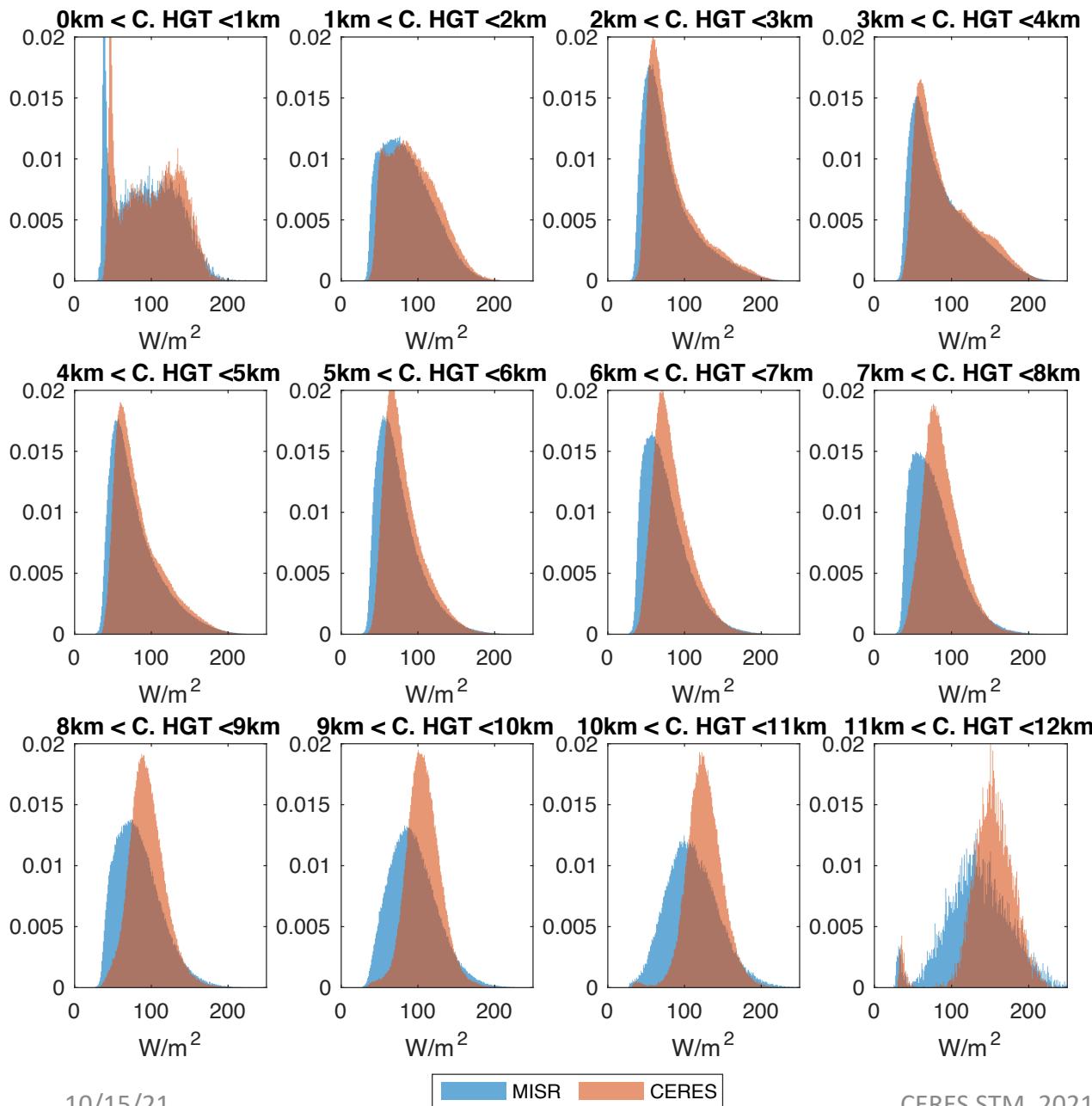
High Elevation

- Different anisotropy?
- 3D effects (Ham et al., 2017)?

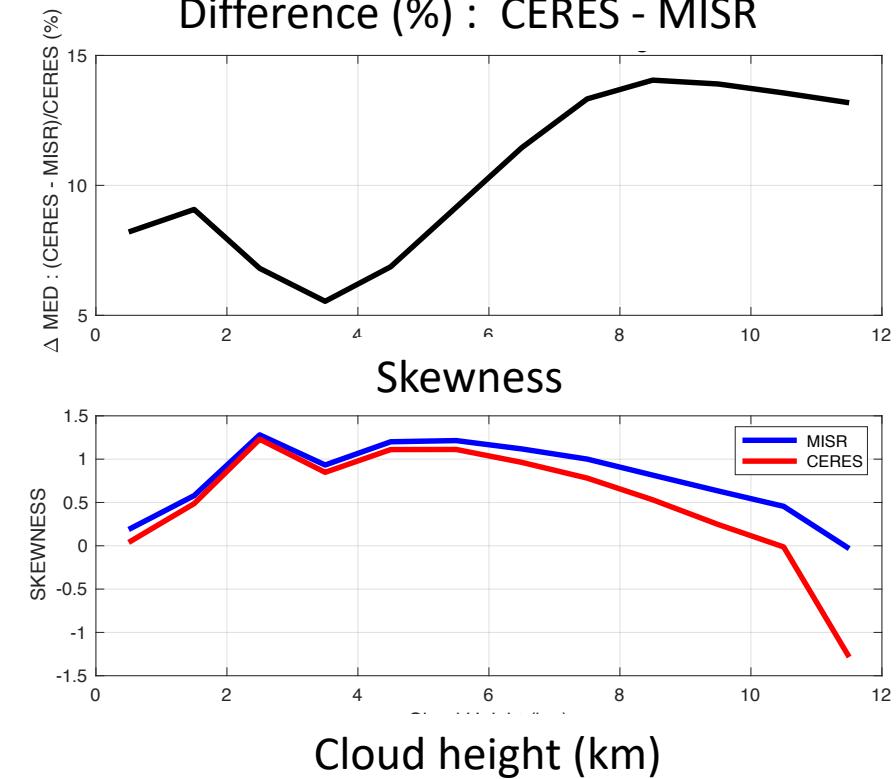


High Clouds

# Cloud Height



Difference (%) : CERES - MISR

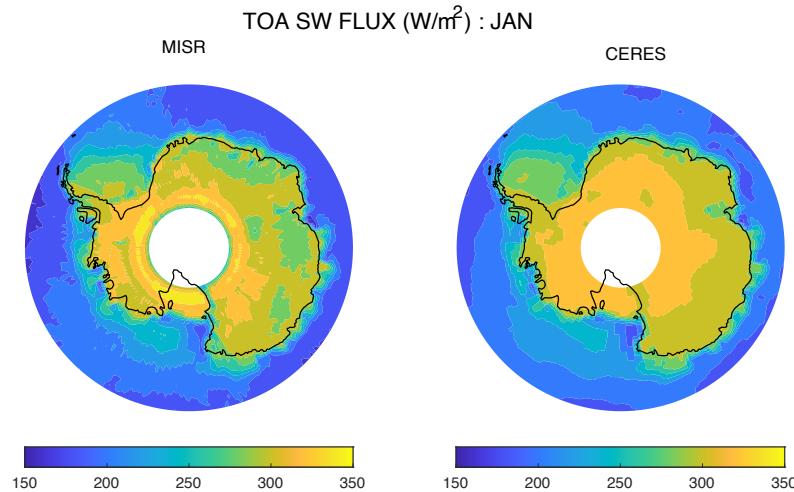




# Antarctic

MODIS, Jan 30, 2009

- Difference between two measurements are outstanding in Antarctica.
  - CERES Flux shows saturated near 350 W/m<sup>2</sup> over Antarctica.

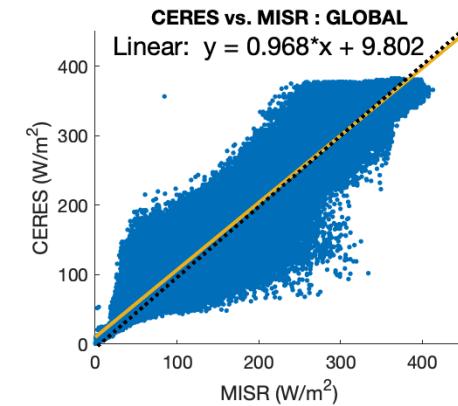


10/15/21

CERES STM, 2021

## 70S-80S : all months

## Global : all months



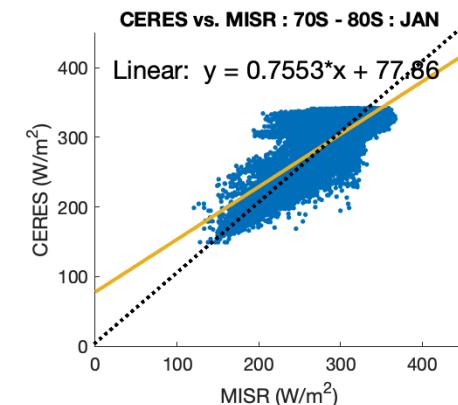
CERES vs. MISR : 70S - 80S : ALL MONTHS

Linear:  $y = 1.004 \cdot x + 0.1252$

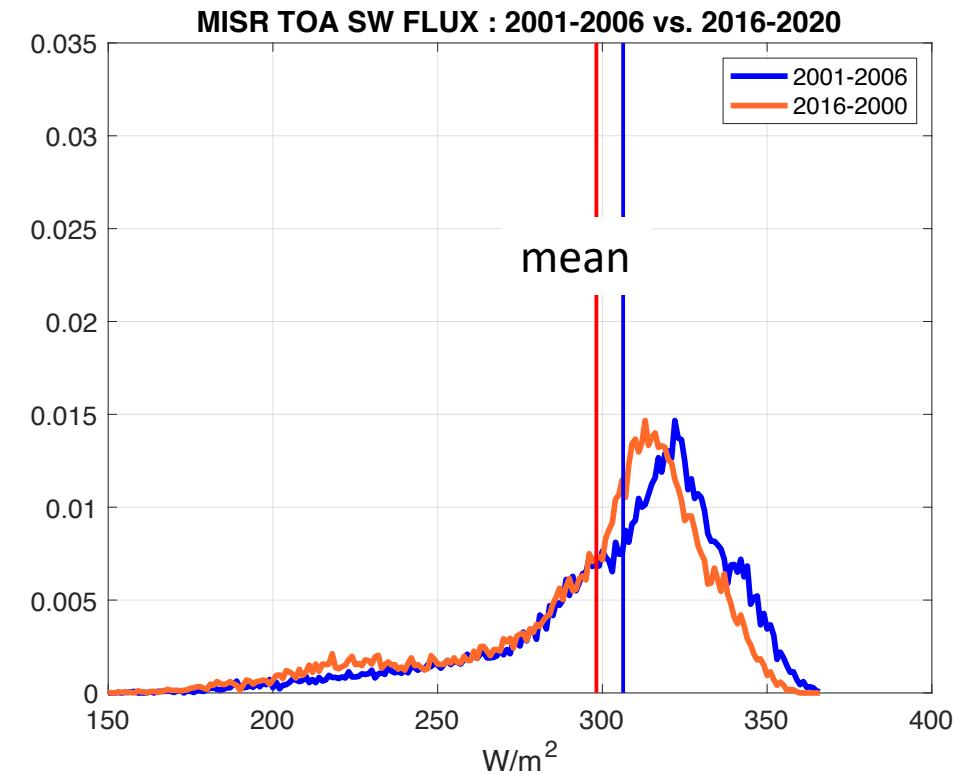
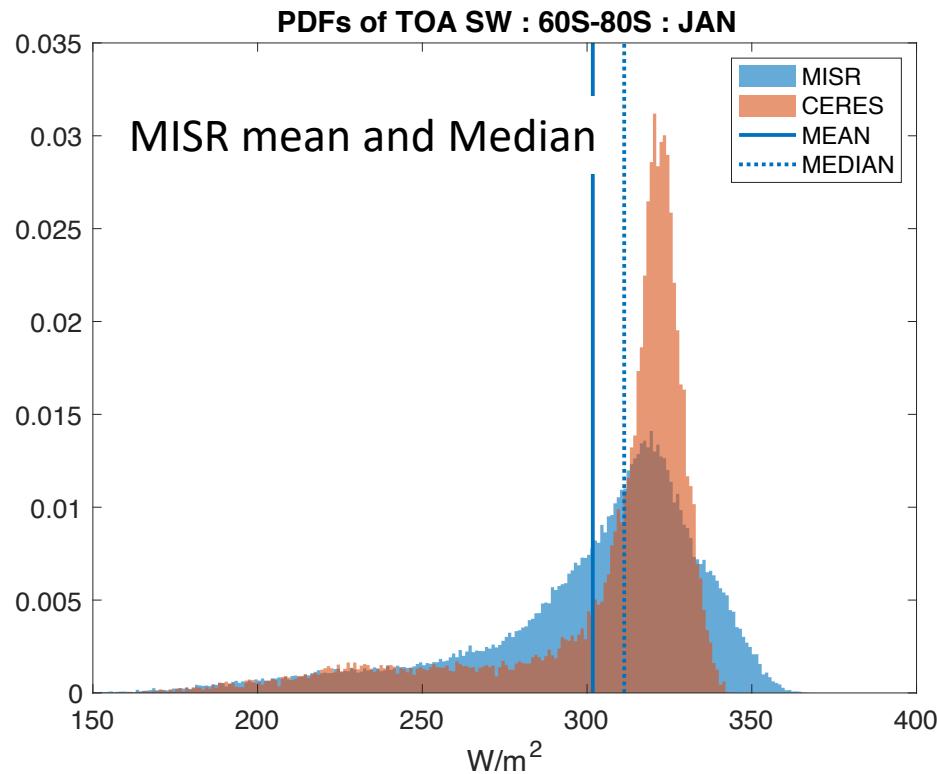
CERES (W/m<sup>2</sup>)

MISR (W/m<sup>2</sup>)

## 70S-80S : January



# Antarctic : 70S-80S : JAN



# summary

- Difference between MISR and CERES SW flux is large at high elevation and high cloud region.
- Should the PDFs be considered with caution in averaging TOA SW flux? Which value can well represent TOA SW flux?
- The PDFs may not be preserved but can change with time and location.
- Potential in MISR spectral albedo retrieval over Arctic and Antarctic, beyond 80S and 80N?